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OFERATION CASTLE---PROJECT 2.2

Report to the Scientific Director

GAMMA RATE VERSUS TIME [4]

(C)by Peter Brown dw Gerald Carp.

U. S. Army Signal Engineering Laboratories Fort Monmouth, N. J.

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ABSTRAC

The objective of Project-2.2 was to measure initial and residual garma rates as a function of time at various distances from high-yield thermonuclear detonations.

Initial-gamma rate versus time was measured at fixed distances from ground zero. In particular, measurements were made of the effect on initial-gamma rate caused by the passage of the shock front from ground zero through the detector station.

Residual-gamma rate versus time was measured to provide gammaradiation time-intensity data, which give information both on fallout rate of arrival and gamma-field-decay rate during the 36-hour period after the defonation.

Scintillation detectors uses used in making measurements. The instrument stations were self-contained -- the only outside facilities required were timing signals to turn on the stations at a predetermined time prior to the detonation.

Data obtained indicate that the expanding fireball and the passage of the shock front from ground zero through the detector station had a marked effect on the initial-gazza rate---hence on the integrated oursure. The initial-gazza rate reached its first peak immediately after the detonation, decreased slowly, begen to rise slowly, and then increased repidly to a second peak (which was about the same value as the first peak). After reaching the second-peak value, the initial-gazza rate decreased rapidly to zero.

The alow decrease in Initial-garma rate was attributed to the natural decay of the fission products; the slow rise, to the expanding of the fireball and approach of the shock front; and the repid increase, to the passage of the shock front through the detector station. These effects were also evidenced in the integrated exposure both prior and subcoquent to the errival of the shock front. The average velocity of the shock front was found to decrease rapidly with distance from ground zero.

The decay exponent from the minidual contraination and follout varied with distance and direction from ground zero, and its absolute value increased rather absorptly several hours after the detonation. This absorpt increase is attributable to the presence of short-lived isotopes in the residual contamination and follout.

The measurements node by Project 2.2 of initial-game rediction exposures from thermonuclear coviers of high yield are in good a warman with data from "The Nuclear Rediction Mendiwok," ANST? 1200, (Selerance 9) (see Figure 4.1).

It eppears that the initial-gooma rediction is of negligible significence, because blast and the seel effects in the seem distance range are so great by comparison that curvival would be possible only if persented were disposed inside blast- and thermal-proof tunkers. ABSTRACT

This report is one of the reports protonting the results of the 34 projects participating in the Military Effects fasts Program of Operation Gastle, which included six test detonations. For medors interested in other pertinent test information, reference is made to UK-934, "Summary Report of the Commander, Tack Unit 13, Programs 1-9," Military Effects Program. This summary report includes the following information of possible general interest: (1) an overall description of each detonation, including yield, height of burst, ground zero location, time of detonation, ambient atmospheric conditions at detonation, etc., for the six abouts; (2) discussion of all project results; (3) a summary of each project, including objectives and results; (4) a couplete listing of all reports devoring the Military Effects Tests Program.

FOREWORD

This report is one of the reports presenting the results of the 34 projects participating in the Military Effects Tests Program of Operation Gastle, which included six test defonations. For residers interested in other partiment test information, reference is made to MR-934, "Summary Report of the Commander, Tack Unit 13, Programs 1-9," Military Effects Program. This summary report includes the following information of possible general interest: (1) an overall description of each detonation, including yield, height of burst, ground care location, time of detonation, ambient atmost plastic conditions at detonation, etc., for the six abots; (2) dissuesion of all project results; (3) a summary of each project, including objectives and results; (4) a complete listing of all reports devoring the Military Effects Testo Program.

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Chopter I INTRODUCTION

1.1 OBJECTIVES

The objectives of Project 2.2 were to measure initial- and residualgamma rates as a function of time at various distances from high-yield thermonuclear detonations.

Prior to Castle, little data had been obtained on the effect of the shock front from high-yield thermonucleur devices on gamma rate and gamma exposure. Project 2.2 was designed to make accurate determination of gamma rates and emposures to be expected at various times and distances from high-yield thermonuclear devices. Comparisons with data obtained from smaller-yield devices may lead to extrapolation factors that would permit more-accurate predictions of reliation levels and decage over a wide range of yields.

Initial-genue rate versus time was determined at various fixed distenses from ground zero; particularly, the offect on initial-genue rate curved by the passage of the shock front from ground zero through the detector station.

Residual-grama rate versus time was measured to provide gavarediction time-intensity data, which give information both on follout rate of marival and gamma-right-docay rate during the 36-hour period efter the detonation.

1.2 PACKGROUND

Productions of the initial-genue rediction from high-yield theorynucleur devices had previously have here a solar plan rade in diver collatively low-yield devices or from from other planes in May data (Previous 1 and 2). Operation Ivy scatter should be tryined and vices did not follow the relativity of the solar har of the world devices. For a vices up to 200 ds, go on radiations the generation of the stars realed the early with yield; for high-yield devices the space of the stars realed the early with yield; for high-yield device a lattice of the stars realed the early with yield; for high-yield device a lattice of the stars of the solar higher on realistic device a lattice of the plan. This we are not an indicated of the stars the star by to the hydrodynesis offers. The planes of the start from the star determines the devices of the marked in a lattice of the planes.

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of the absorption medium, and produced a consequent change in the shape of the gazza-rate-versus-time curves during this period.

1.3 THEORY

The gamma radiation emitted from a nuclear detonation may be divided into two portions: initial and residual rediation. The residual portion may include radiation both from fallout and from neutron-induced activity.

1.3.1 Initial-Garma Rediation. For devices with yields of less than 100 kt, in which the hydrodynamic effect is small, the initial rediations are divided approximately as shown in Table 1.1 (from Seference 4). The

TABLE 1.1 FREEDOY PRATITICS ANONG INITIAL PROTATIONS FOR FISSION-TYPE DEVICES

Pros Beference 4

والمحتليسين بالمتثلثة بالألبان ستعمد والوجادي والوائدة متالية المراجع ومتراديا	ومربيكات فكرعة بالتكريبة فبالمناطقة والمتحديناتها	and the second sec
Pechanian	Percant of Total Fission Savrgy	Total Scorgy For Sizaton
		2014
Froaps Houtrons	4.0	8
Pro at Casas	4.0	8
Maslan Disclose Gumas	2.7	5.4
Pivalon Prolant Lates	2.7	5.4
Fission Product Heutrinos	5.5	ц
Dultysd Hautrops	0.1	0.2

"Hostly superhed in the device

major contributions to initial-gamma rediction are from the fission-product gamma and from the gamma rediction from nutron capture by R^{14} (n,y) in the high-simploaive components and in air. The prospt gammas are nearly all absorbed in the device itself and are of little rightflernes outside the device. The fiscion-product gammes productionts at close distances (Reference 4). The R^{14} (n,y) grains between increasingly important at greator distances, and eventually become the rejer contributor. Figure 1.1 shows the contribution from fiscion-product gammes and R^{14} (n,y) for a light carfore burst. With respect to the n R^{14} (n,y) modulation is even tiply soluted within 0.2 meanly the fiscion-product gamme, lowever, continue to contribute for the fight 20 module.

For this sublest divides, it is providery to evolve the indice the indication of next and find to its provident billing in the little to provide the test from finders provide general. Show it then the to the indicate general any very over this little, drive they on it is found as the draw it is a given yield, the prover of centrative distingtion of the test is a grant fundom as for filmion, and therefore a long or shift is in the grant of the tion of for filmion, and therefore a long or shift is to grant of the tion of for filmion, and therefore a long or shift is a long or of the tion experies in the to the M^4 (a.g) is obtained as a therefore during the vice (defermend).

1.3.2 Residualed on a Redition. The residual-good of distinguent and in terms of a sector of the following the sector of the sec

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512.37

induced activity. The decay rate of the residual radiation from fallout will follow approximately the expression:

and

$$r = \int_{t_1}^{t_2} I_t \, dt = 5I_1 \, (t_1^{-0.2} - t_2^{-0.2})$$
 (1.1)

Where: I_t = exposure rate at time t

 $I_1 = exposure rate at unit time$

t = time

 $\mathbf{r} = \exp \operatorname{posure}$ between times $\mathbf{t_1}$ and $\mathbf{t_2}$, where $\mathbf{t_1} \ge 10$ seconds

It is expected that the decay of the residual radiation will vary wit weapon design. For example, the presence of Np³⁰ would tend to decrease the absolute value of the decay exponent for a period of time.

1.3.3 Absorption in Air. The absorption of unceattered gamma radiation in homogeneous air is exponential with distance. From a point source of monophyrgatic radiation, the variation of intensity with distance is expressed as:

(1.2)

$$I_{D} = \frac{I_0 \exp(-\mu D)}{4\pi D^2}$$

Where: I_{D} = intensity at distance D

- I₀ = source intensity
- = total linear abcorption doefficient (this
 coefficient goverally decreases with
 Increasing gamma energy)
- D = distance

The absorption coefficient μ in Equation 1.2 is applied by a waveterm generative, and a contraction should be only for field coefficient where the detector is equations for a function point, that is does by elding a building factor B to Constant 1.2 to second for the continued ondiction that will be detected. Initial-up forther for distance exercises

11

SCORT

and distances have been calculated (Reference 5), and some values are shown in Table 1.2. For omnidirectional detectors, the expression is:

$$I_{D} = \frac{I_{0}B \exp(-\mu D)}{4\pi D^{2}}$$
(1.3)

1.3.4 Hydrodynamic Effect. As shown in Section 1.3.3, the attenuation of gamma radiation is highly dependent on the amount of absorber between the source and the detector. For devices of less than 100-kt



Figure 1.1 Gazza exposure variated distance for a 1-bt surface burst. This illustration above the contribution from fiesdon-product general and k^{14} (n,y).

yield, essentially all of the initial-gener rediation is emitted hofore the check front can produce an approxiable of age in the effective electrotion of the nir between source and detector. For high-yield devices, the velocity of the check front in proficiently high to produce a strong on homoment of a large percentage of the initial-general relation (Palarence

12

6). The higher the yield, the larger is this percentage. A simplified treatment of the hydrodynamic effect follows.

Assuming a sphere of volume V_0 , radius R, and filled with a gas of density ρ_0 and mass M, then

$$M = V_{0}\rho_{0} = \frac{4\pi R^{3}\rho_{0}}{3}$$
 (1.4)

Let the gas be compressed into a shell with thickness ΔR (R remaining constant). The new gas volume is expressed as V_1 with a density of ρ_0 ($V_1 = 4\pi R^2 \Delta R$). The mass has not changed; thus

$$M = V_{c}\rho_{0} \doteq 4\pi R^{2} \Delta R\rho_{1} \qquad (\Delta R << R)$$

$$\frac{4\pi R^{3}\rho}{3} \doteq 4\pi R^{2} \Delta R\rho_{1} \qquad (1.5)$$

$$\Delta R \rho_i \doteq \frac{R \rho_0}{3} \tag{1.6}$$

Equation 1.6 indicates that a ray originating in the center of the sphere would traverse only a third as such of the mass in the shell model

TABLE 2.2 CALCULATED BUILD-UP FACTORS

The build-up factor (B) given here is the factor $B_{\rm F}$ ($\mu_0 D_{\rm s}, E_0$) as computed by the Muchaar Development Associates for AFSWF (Reference 5).

Emergy (E,)	Bu	Lid-up Factor,	8
	1,000 yards	1,500 yeads	J COO YATE
Hav			
1	16.2	29.3	~95
3	3.85	5.35	10.2
4	2.97	4.00	7.00
10	1.70	2.01	2.90

as it would in the homogeneous model. The result would be an enhancement of radiation. Once the shell of material in the shock front purses the detector, an over-greater enhancement should result.

As proviously stated, the $h^{14}(n, \gamma)$ corporant of initial radiation is essentially emitted within 0.2 second. Since it takes at least one escord for the shock front to reach a detector st a distance of 7,000 feet (even for devices in the order of 6 ht), the $h^{14}(n, \gamma)$ corporant is not significantly enhanced. The fitsion-product general continue to contribute during the first 30 seconds; therefore, this radiation is strongly enhanced by the shock wave.

Chapter 2 PROCEDUR<mark>E</mark>

2.1 INSTRUMENTATION

A low-sensitivity system was used for measuring high-initial- and high-residual-gamma radiation rates. For measuring lower residual-gamma rediation rates from fallout, the instrumentation was of high sensitivity. (See Appendix for circuit theory of the instrumentation.)

2.1.1 Low-Sensitivity System. The detecting element was a stilbene crystal mounted in a graphita block for electron equilibrium. This crystal assembly was installed inside a blast-resistant housing at the top of a light pipe. The crystal output passed through the light pipe and was detected by a IP21 photomultiplier tube. This tube was used in a 100percent feed-back circuit, which held the anode current nearly constant (regardless of the incident light) by reducing the photomultiplier-tube dynode voltage. The gain of the tube was approximately a direct function of the antilog of the dynode voltage. In this manner, a useful dynamic range of about 10 was realized.

Two different recording systems were used to take edvantage of the wide dynamic range of the instrument. Initial-gamma rates from 10³ r/hr to 10⁹ r/hr were recorded by a high-speed recorder, which was capable of resolving to 0.01 second with a recording time to 20 minutes. This recorder, manufactured by the Samborn Company, Cambridge, Massachusetts, used beat-sensitive paper and a hot-wire stylus. It was found to be wellsuited to recording under the adverse conditions encountored during nuclear tests, since it was virtually free of the troubles that may arise with ink and photographic recorders.

Initial-gauma rates from 0.20 r/hr to 10^4 r/hr with low resolution over a long-time period were recorded by a Bristol recording millivoltmater. This recorder used synked paper, had a running time of about 24 hours, and a time resolution that was proportional to the running time (1 minute at the start and 5 minutes at the end of 24 hours).

A double station, consisting of two detectors, one with a low- and one with a high-resolution recording system, covered the range from 0.2 r/hr to 10⁹ r/hr and recorded upeful data from 0.01 second to 24 hours.

2.1.2 High-Sonsitivity System. The detecting obsent was an anthracone crystal mounted in a graphice block for electron equilibrium and attached to the window of a 5019 photomultiplier tube. This tube, like the IP21 mentioned in Section 2.1.1, was used in a 100-percent feed-back circuit, and the encde current was kept virtually constant. Its gain was approximately a direct function of the antilog of the dynode voltage.

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Therefore, it was possible to plot the incident light, or gamma radiation, as a function of the dynode voltage.

All electronics were mounted inside a light-aluminum housing for protection from weather, thermal influences, and small blast pressures. The output was recorded on a Bristol recording millivoltater, which recorded gamma rates from 0.2 r/hr to 10⁴ r/hr for a period from 1 minute to 24 hours.

For field installations, the instrumentation systems described in Sections 2.1.1 and 2.1.2 were mounted inside standard 55-gallon stuel druns imbedded in concrete. Figure 2.1 is a block diagram of the detector systems.

2.1.3 Special Detector System. A special detector system was designed and constructed to determine the effects of high-gamma rates and



Figure 2.1 Block diagrees of a typical detector system. The detecting element is of ther a stillens or an enthracian crystal depending on whether it is part of a low- or a high-sensitivity system.

cumulative exposures on the crystal detecting element. Two low-constituity detector heads were employed: one in normal usage and the second covered with a 5-inch-thick layer of lead shot to attenuate the garma rediation reaching the crystal by a factor of about 100. The "normal" head detected

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the incident gamma rate, while the lead-covered one detected 1/100 of this rate, and also received only 1/100 of the total exposure.

To calibrate the station with the normal head, the first method was used as described in Soction 2.2.1. This required the placement of a density-four neutral filter between the crystal and the photomultiplier tube of the normal station to increase the range to 10⁸ r/hr. The leadcovered head required only a density-two neutral filter, inasmuch as the lead covering served to attenuate the incident gamma radiation by a factor of 100 (equivalent to a density-two filter).

If the two delector heads track, the crystals are not being differentially affected by gamma rates and cumulative exposures. If the detectors do not track, it is an indication that there is a rate or cumulative exposure dependence of the system or a significant change in the spectral quality of incident radiation. In the event that scintillation crystals are shown to be neither rate nor desage dependent, the assumption may be made that a nearly linear extrapolation of low-intensity calibration can be performed.

2.2 CALIBRATION

2.2.1 <u>High-Intensity Calibration</u>. Calibration for high rates was based on the assumption that the light output of the scintillation crystal is a linear function of the incident radiation at least up to 10^9 r/hr.

Two techniques were used to calibrate the high-level instruments. In the first method, the detector head was calibrated from 0.2 r/hr to 10^4 r/hr, using an 88-curie Co⁶⁰ source. To extend the range of the instrument to 10^3 r/hr, a density-four neutral filter was placed between the crystal and the photomultiplier tube. Effectively, the system would then measure radiation from about 0.2 x 10^6 r/hr to 10^6 r/hr.

It was necessary to put chielding around the photomultiplier tube, since it was sensitive to rediation. The purpose of the shielding was to attempt to attenuete the incident gamma rediation on the photomultiplies by the same order of magnitude as the noutral density filter attenuates the light from the crystel. (This was not possible in the field, as the amount of scatter radiation striking the photomultiplier through the light pipe procluded an attenuation factor greater than 10¹.)

In the escend high-intensity calibration method, the detector head was first calibrated from 0.2 r/hr to 10^4 r/hr using an 85-curie Co⁶⁰ source. The crystal was then replaced by a light cource of the same spectral quality as the stilbane crystal output. The output of this light source could be varied in known increments. A plot was made of this output versus the relative incident luminous flux. The coldbration curve thus obtained was then normalized to the curve obtained with radiation and was found to track very clocely up to the meximum of 10^4 r/hr (Figure 2.2). Essence of the egreement between the light and the rediation calibrations at the low levels, the light calibration was used as a basis for extrapolating the radiation curves up to 10^3 r/hr.





TABLE 2.1 INSTRUMENTATION FOR EACH SHOT

Station Number	Island Location	Raige from Zero	De scription.
		1002	
Shot 21			
220.08	Obce	83.762	Single high-monaitivity unit
220-10	Able	6,400	Two low-sensitivity units
المد تصد	Consile	7,86	The low-sensitivity units
		-	Two low-sensitivity units with
220.12	Dog	41,372	one datector med shielded with Po
220.13	- Luy	45,279	Single high-mentivity unit
220.14	Peter	82,645	Single high-masitivity unit
221.01	William	61,719	Single high-monstivity unit
221.02	Yoke	54,480	Single high-consistivity unit
221.03	Zebre	50,598	Single high-sensitivity unit
221.04	ALFA	49,432	Single high-sensitivity unit
221.05	Brevo	47,590	Single high-monaitivity unit
Shot 21			
220.08	0638	~ 83,800	Single high-seasitivity unit
220.12	Dog	~41,400	Sincle high-senditivity unit
220.14	Peter	~ 82,700	Sincle algureinsitivity unit
Shot Ji			
220.08	0500	26,778	One low- and one high-mensitivity unit
220.09	Roger	7,511	One low and one high sensitivity unit
270.12	Dog	69,112	Single bigh-menal livity whit
220.14	Poter	11,779	One low and one high-pensitivity unit
Sbot 41			
220.06	Tox	13,501	Ope low- and one high-sensitivity unit
220.07	George	15.579	One low and one high-sensitivity unit
220.12	Dog	7,17	Typ low-suncitivity units
220.13	Reay	9,602	Cos bigh-monsitivity unit
Shot 51			
Уола			
Shot 61			
Vanuabered	Allos	~18,000	Two portable high-monsitivity units

s,

2.2.2 Low-Intensity Collibration. The low-level detectors for measuring fallout wore collibrated directly with an 85-curio Co⁶⁰ course.

2.2.3 Energy Dyrandance. Both the anthrasene and stillers detector systems gave approximiting inter response over the every were from 100 key to 5 Nev. The pertoperation laboratory recalibration of the anthratere and stillers detectors that was originally placed could not be dore.

2.3 SHOT PARTICIPATION

Table 2.1 summarizes the location, range, and types of instrumentation used in each of the shots.

18

Station	1)pe	Distance to	Asimuth		Cares IX	-mroo		Integrated	Yaxima Kito	Decay Exponent	
		ater purch		Frior to of Shock	Arrival Front	Subsequent of Shock I	to Arrival				1
		Fut	de L min	-	Percent of Integrated	4	Parcent of Integrated		r∕br		
Shot Li											
220.11 Charles	High Intensity	7,306	8) X	9,900	1	1	1	9,900 (0.9 me)	or = 9	i	* 3 5 7 5
0000 5.2.022	Fallout	83,762	ររូ៖ ស	ł	I	1	1	133 (0-25 6 20 年)	¥	(1 0.61 18 LT)	
2 E E	Pallout	uc'n	22 24	l	1	1	ł	3,735 (0.25 to 20 hr)	1,000	11 12 12 12 12 12 12 12 13 13 13 13 13 13 13 13 13 13 13 13 13	
11 1002	Fullout	6) 111	89 24	1	;	ł	ł	12 12	4	1.24	
				•	forment of	4	Report of				
Shot 41		Ĩ		~	Parcent of Integrated	۳	Narcant of Integrated		7/14		
223.12 Dog	ligh Intensity	מנינ	× ~	7.2 (0.62 w 1.3 m		12,121 (1.) (ة. ع ك ك	13,163 (0.02 to 20 me)	7.2 x 10 ⁴	1	F S S S E
2.20 .06 Fox	kigh fatansity	10,501	и и	10.5 {0.02 to 1.5 =	تع ۲	25.5	3 3	91.2 (0.42 to 25 me)	5.2 x 10 ⁶	ł	1953E
220.37 Georys	Pailout	15,579	R 13	ţ	ł	1	1	1,237 (Ano) (0.031 to 10 hr)	(Ar a)	(1 07 3 2) 5 3 2) 5 4 9 5 1 2)	11887 F

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decay exponent x, in the expression $I_t = I_1 t^{-x}$, was found to change from about 1.35 during the period H + 3 to H + 10 hours to about 2.1 during the next fourteen hours.

Figure 3.2 shows that at Station 220.08, fallout started to build up about 15 minutes after the detonation and reached a maximum rate of 34 r/hr. The integrated exposure was 133 r during the first 24 hours. The decay exponent was 0.81 for the period H + 1 to about H + 18 hours.

Because of the unexpected high yield of Shot 1, the instrumentation sustained serious damage. Low-sensitivity Station 220.10 on Able was destroyed by blast. Fallout Stations 221.05 on Bravo, 221.04 on Alfa,



Figure 3.1 Initial-gamma rate versus time for Shot 1, Station 220.11 on Charlie. This station was 7,206 feet from ground zero.

221.03 on Zebra, 221.02 on Yoks, and 221.01 on William either Hure struck by missiles or inundated by water. There stations did not provide reliable records.

3.1.2 Shot 2. Instrument stations at Able and Charlis were destroyed by high overplassures from Shot 1; therefore, Shot 2 could not be instrumented for close-in initial-garma-rate masurements. Fallout from

21





Plure 3.4 Lattial-gumm rate wreat the for Sot 4, Station 220,12 on Det. This station was 7,12 but from ground arts.



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Figure 3.5 Tertial-goume rate versus then for Shot 4, Station 220.06 on Post- Tais station was 13, 501 foot from ground mark.

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Shot 2 was of such a low level compared to existing background that fallout stations on Obos, Dog, and Peter did not obtain good records.

3.1.3 Shot 3. No initial-gamma rate data were obtained because of the unexpected low yield of Shot 3. Fallout Station 220.12 on Dog provided a good record (see Figure 3.3). Fallout started to build up about 20 minutes after the detonation and reached a maximum of 24 r/hr. Integrated exposure was 51 r in the first 15 hours. The decay exponent was 1.28 for the period H + 1 to H + 12 hours.

3.1.4 Shot 4. Figure 3.4 (Station 220.12 on Dog) shows that the initial-gamma rate first dropped in magnitude, then increased alowly from 0.4 to about 1.2 seconds, and then increased relatively rapidly to a value of 7.2 x 10^6 r/hr, approximately the same as the original rate at time of detonation. The slow increase appears to be the combined result of the expension of the fireball and the approach of the shock front toward the detector station.

Figure 3.5 (Station 220.06 on Fox) shows the same phonomena as Figure 3.4. The expanding fireball and the approach of the shock front cause a slow increase in the gamma rate from 0.7 to 4.48 seconds. Passage of the shock front through the detector station causes an abrupt increase to about 4.5×10^3 r/hr, after 7 seconds. This peak is about the same as the original rate at time of detonation.

The record obtained from Station 220.07 on George shows an initialgamma spike and the arrival of the shock front 5.86 seconds later. However, the deflection was too small to permit the readout of a time history of the initial gamma. From Shot 4 data, the average volocities of the shock front between ground zero and Dog, Fox, and George are computed to be about 5,100, 3,000, and 2,600 ft/sec, respectively.

Finidual-genua rate versus time is plotted in Figure 3.6 (Station 220.07 on George). Fallout reached a peak value of 620 r/hr, and the recorded exposure of the first 10 hours after detonation was 1,237 r. During the period H + 2 to H + 6 hours, the decay exponent was 1.39, and from H + 7 to H + 10 hours it was 1.61.

These figures represent the data as read; no correction for air density has been avela. The relative air density for the Castle events is 0.9.

3.2 DATA CORPELATION

3.2.1 Integrated Exposure Pate Versus Total Exposure. The decegos obtained by Integrating the curves of Figures 3.2 and 3.6 have been compared to decays measurements made by the National Diracu of Standards (NBS) type photographic decimaters on Project 2.1 at the same station locations (Reference 8). The results are gumenized in Table 3.2. Because the genue rates obtained are valid at a sufficient number of mediumdistance stations where exposure did not exceed 10°r, it may be assumed that the results are also valid at close-in stations where the total ex-



Figure 3.6 Besidual-gamma rate versue time for Shot 4, Station 220.07 on George. This station was 15,5-9 feet from ground serve



posure was less than 10^8 r. This assumption is based on data showing that the response of scintillation-type crystals is not adversely affected by total exposures of less than 10^8 r (Reference 7).

3.2.2 Special Detector System. The data from Station 220.11 on Charlie, Shot I, showed that the two heads did not track. This was attributed to the presence of scattered radiation that entered the shield-

TABLE 3.2 - GANNA RADIATION DUSAGE

Comparison of desage measurements made with NES type photographic desimators on Project 2.1 and desage measurements at sems ranges made by solutillating detectors on Project 2.2.

Distance to Ground Zero	Project	Station	Dope	Time
feet			F	br
Shot 1:				
83,762	2.1 2.2	210.2 3 220.08	180 133	н + 82 н + 20
41,372	2.1 2.2	210 .12 220 .12	6,000 3,735	н + 78 н + 20
Shot 41				
15,579	2.1 2.2	210.21 220.07	2,060 1,237	H + 104 H + 10

ed area through the light pipe. The consitivity of the photomultiplier tube to radiation required the presence of shielding to attenuate the radiation in the same order of magnitude as the neutral-density filter. When it proved impractical to increase the shielding, the use of neutraldensity filters to increase the detector's upper limit was abandoned for remaining shots.

Chopter 4 CONCLUSIONS

The following conclusions were reached, based on the empirical results of Project 2.2, and on collateral data obtained from other Castle projects:

The passage of the shock front from ground zero to and through the detector station has a marked effect on the initial-gamma radiation rate for high-yield weapons. The percentage of total-initial-gamma-radiation exposure received prior and subsequent to the arrival of the shock front was found to be a function of the distances from ground zero.

Initial-gamma radiation is of negligible significance since blast and thermal effects in the same range of distances are so great that survival would be possible only if personnel were disposed inside blastand thermal-proof bunkers.

The rate of decay of the residual field radiation from fallout was found to vary with distance and direction from ground zero. The rate of decay increases rather abruptly several hours after the detonation. This may be attributed to the presence of short-lived isotopes in the residual contamination and to iractionalization of fallout materials.

Figure 4.1 indicates that Project 2.2 measurements of initial-gamma radiation exposures from thermonuclear devices of high yield are in good agreement with data from "The Nuclear Endiation Handbook," AFSWP 1100. It should be noted, however, that the Castle Project 2.2 data was one of the sources used in the preparation of the prediction curves of that handbook.

Chapter 3 RESULTS and DISCUSSION

3.1 CENERAL

Figures 3.1 to 3.6 are plots of gamma rate versus time at various distances from ground zero.

Examination of the figures makes apparent the effect of the shockfront passage between ground zero and the detector station during the initial-gamma radiation. After the passage of the shock front, the initialgamma rate rose almost to the same value as it was at the time of detonation.

Table 3.1 summarizes the gazma-radiation data for the stations used during Shots 1, 3, and 4. The extensive instrumentation originally planned was not followed because of the unexpected loss of most of the equipment stored on Table during Shot 1. Out of a total of 30 low- and 5 highsensitivity instruments, only seven of the former and two of the latter were salvaged for succeeding shots.

3.1.1 Shot 1. Figure 3.1 shows initial-gamma rate versus time for Station 2.0.11 on Charlie. The increase in rate from 0.35 to 0.9 second was apparently caused by the expanding fireball and the approach of the shock front toward the detector station. The shock front arrived at the detector station at 0.9 second. The average velocity of the shock front between ground zero and the detector station was computed to be about 8,000 ft/sec. No further data were obtained because the station was destroyed by missiles at that time.

The special detector system at Station 220.11 on Charlie was designed to datamino the rate or cumulative exposure dependence of the crystal dotecting elements, as well as to measure the initial-gamma rate. Records obtained showed that the two detector heads did not track. However, it was determined that this was not due to a rate or deaved dependence of the crystals, but rather to scattered gama radiation entering through the light pipe and striking the photomultiplior tube in the station without a shield around the detector head. Since a density-four neutral filter was placed between the crystals and the photomaltiplier tube to increase the range to 10° r/hr, the effect of the scattered rediction entering through the light pipe was to multiply the direct effect of games radiation on the photomeultipliar by reveral orders of regnitude. Economy of this, the record wont off scale is collaiely and gave no true indication of the games rate. The station with the shielded detector head provided a good record, because the lead chielding effectively attenuated the game rediction that reached the phototube by a factor of 100.

Figure 3.2 plots residual-gauge rate versus time for Stations 200.12 and 220.08. The graph shows that at Station 220.12, follows started to build up 15 minutes after the detonation and reached a minimum rate of 10³ r/hr. The integrated exposure for the first 24 hours was 3,735 r. The

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Appendix

THEORY OF INSTRUMENT OPERATION

Buffiliment of the objectives of this project required an instrument of known operational charasteristics in vary-high radiation fields, and with an extremely large dynamic range. The former requirement indicated that a scinvillation detector would be tost suited; the latter led to the choice of a feed-back circuit that kept the photomultipliar-tube anode current substantially constant and thus prevented fatigue of the photoesthede under high flux.

The photomultipliar dynods voltage is developed across the cathods resistor of a cathods follower, whose input signal is derived from the photomultipliar mode current (Figure A.1). in increase in the incident luminour flux on the photocathode tends to increase the photomultipliar-



Figure All Schoolstic discrets of a typical detector system. High-intensity measurements were made with a 1921 photomultiplier tube whereas fallout measurements were made with a 5819-type tube.

tube anothe current. Increased anois current presents a noisbive signal to the control tube, incaracing the impedence of that tube and loweling the solice a corner the photonultiplicarities dyongles. Since the emploid and the photonoltiplier tube is a function of dynode voltage, the lowest dynode voltage will cause a decomposition of a context. An equilibrium condition is previous at a point where a shell increase is photonoltiplier and on our out decreases the dynode state of the low of the tube to a value that is just enough to keep the could durant at the new low lot biper-incident light flux. The output is taken from the cutode of the control: tube end for the tube to a value the taken from the cutode of the control:

The range of ano becorrect explation can be extrulated for a given control tube and foodback resistor. If the control tube has a grid back of 5 volts (that is, the tube cuts off at

20

-5 volts), and the feed-back relator is 100 megohas, the maximum value of the such current $I_{\rm m}$ is given by:

 $I_{g} = \frac{E}{R} = \frac{5}{10^{2}} = 5 \times 10^{-3} \text{ umperes } (0.05 \, \mu \text{s})$

Thus, in a typical circuit the photomultiplier-tube anode current cannot exceed 0.05 µe, and at this value of current there is no danger of photocethode fatigue. Since this was a 100-percent feed-back circuit, it was wirtually independent of control-tube characteristics and depended only on the photomultiplier-tube characteristic and the value of the feed-back resistor. The response curve of the instrument closely followed the photomultiplier tube's gain characteristic for warying dynode voltage.

The tube complement included a 104 as control, a 1921 photocultiplier in the low-sensitivity (high-intensity measurements) cycles, and a S019 photocultiplier for fallout measurements (highscientivity system). Although the fated plate-to-filement voltage of the 104 was encoded by a factor of ten, no difficulties were experienced if filements were allowed to ware up before applying the plate voltage. A 45-wolt bettery applied a constant voltage tetween anode and minth dynote of the photominiciplier, since the latter is most stable when operated with a constant voltage between these points.

No variables in instrument characteristics could be detected when the control tubes were replaced. This was true even when tubes you triad with different characteristics (such as the 134) but similar pin communitors. This substantiates the theory that in this 100-percent feedback circuit, the plotomultiplier-tube characteristic is independent of the control-tube parameters. The dynamic range of the instrument is approximately 10°. When the extreme value of dynamic range is not required, the recorder may be edjusted to operate over a portion of the range, thus providing an expanded scale over the region of interest. For example, in some locations where a dynamic range of 10 was sufficient, the recorder was edjusted to operate over this portion of the instrument's output. A more-accurate recording was provided by making these sijustants.

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